

Paschen¹ that if the vapour pressure in his lamp was increased, the relative intensity of the line 1.014μ came out still higher, while with a low vapour-pressure the intensity of the two lines was about equal. This may explain the values of the intensities found for these lines in the present investigation.

Table II.

Wave-length.	Frequency	Difference.
First subordinate triplet series. $r = 2, p-m, d, m = 3.$		
3663.05 A° U.	27292.06	4630.92
3131.66 " "	31922.98	1767.36
2967.37 " "	33690.34	
Second subordinate triplet series. $= 2, p-m, s, m = 1.5.$		
5460.97 A° U.	18306.73	4631.31
4358.66 " "	22937.04	1767.19
4046.78 " "	24707.23	
Triplet 1.038μ 1.270μ 3.02μ .		
30200 A° U.	3310.3	4561.5
12700 " "	7871.8	1759.6
10380 " "	9631.4	
Suggested triplet 1.09μ 1.367μ 3.70μ		
37000 A° U.	2702.7	4608.6
13670 " "	7311.3	1860.4
10900 " "	9171.7	

In looking for series relationships among the lines given in the first column of Table I., it was seen that the frequency differences for the lines 1.038μ , 1.27μ and 3.02μ are practically the same as those which characterise the subordinate series triplets in the mercury spectrum given by $\gamma = 2, p-m, d$ and $\gamma = 2, p-m, s$. This will be evident from the numbers given in Table II. It will be noted, too, that the frequency difference between the line given in our list at 1.09μ and the one given by Paschen at 1.367 is equal to 1860.4 , which approximates, as the table shows, to the frequency difference between the second and third numbers of the triplets of the two subordinate series mentioned above. If these two lines should turn out to be the second and third numbers of a triplet similar to the one

¹Paschen, Ann. d. Phys. 27. 13. p. 559, 1908.