The main change is in the eccentricity, which was greatly over-estimated in the preliminary elements. In proceeding to a second solution ω was put at 90 degrees, and a small change made in T to correspond. The eccentricity was made 0.05, and the solution again carried through. The normal equations using Schlesinger's notation are:

$$\begin{array}{l} 10 \cdot 200\tau + 0 \cdot 808\kappa - 1 \cdot 249\pi + 0 \cdot 705\epsilon - 0 \cdot 652\tau = -1 \cdot 586 \\ + 4 \cdot 863\kappa - 0 \cdot 774\pi + 0 \cdot 299\epsilon - 0 \cdot 719\tau = +1 \cdot 114 \\ + 5 \cdot 337\pi + 0 \cdot 182\epsilon + 4 \cdot 765\tau = +0 \cdot 352 \\ + 1 \cdot 300\epsilon + 0 \cdot 209\tau = +0 \cdot 425 \\ + 4 \cdot 285\tau = +0 \cdot 190 \end{array}$$

whence, $\tau = -5.54$ dT = -.125 $\epsilon = +0.44$ de = -.008 $\tau = +5.037$ $d\omega = -11.73$ $\kappa = +0.174$ dK = +.174r = +0.063 $d\gamma = -.19$

and the final elements with the probable errors become,

$$\begin{array}{rcl} P & = 3.854 \; \mathrm{days} \\ e & = 0.042 & \pm .025 \\ T & = \mathrm{J.~D.~2,420,370 \cdot 259} & \pm .351 \; \mathrm{day} \\ \omega & = 78^{\circ} \cdot 27 & \pm .21^{\circ} \cdot 9 \\ K & = 24 \cdot 77 \; \mathrm{km.} & \pm .70 \; \mathrm{km.} \\ \gamma & = +7 \cdot 81 \; \mathrm{km.} \\ a \; \mathrm{sin} \; i & = 1,312,000 \; \mathrm{km.} \\ \frac{m_i^3 \; \mathrm{sin}^3 \; i}{(m+m_i)^2} & = 0.0061 \odot \end{array}$$

The residuals from these elements agree with the residuals from the observation equations to the nearest tenth of a kilometre. Comparison with the residuals given by the first solution shows that little has been gained, and the differences between the last two sets of elements are less than the probable errors.

The residuals given by the individual plates are tabulated under the heading (O-C) in the table of observations. The probable error of a single observation is 3.7 km.

Dominion Observatory Ottawa May, 1917.