

MEAN DISTANCE OF STARS WHOSE RADIAL VELOCITIES, PROPER
MOTIONS AND PARALLAXES HAVE BEEN
DETERMINED.

BY REYNOLD K. YOUNG, Ph.D.

The following investigation was undertaken to answer the question,—
Will the mean distance of the stars as determined from their proper motions
and radial velocities agree with the mean distance as formed from the
directly measured parallaxes?

The data for the parallaxes were selected from the list published by
Kapteyn and Weersma in 1910*. For one hundred and ninety-five stars
of this list, radial velocities were obtained from the Mount Wilson† and Lick
Observatory‡ results. The proper motions of all these stars are very large,
and indeed for the most part the stars were selected for parallax measure-
ment on this account. The excessive magnitude of the cross motion may
be due

1. to the proximity of the stars,
2. to the excessive speed,
3. to exceptionally large values of the inclination of the motion to
the line of sight.

An examination of the data seemed to show that the first factor was
by far the most potent. If we reject twenty-eight stars with exceptional
velocities, over fifty kilometres per second, the remaining ones have a mean
radial velocity very little above the average of the stars in general.

*Gronigen Publications, No. 21.

†Ap. J. Vol. 39.

‡Lick Observatory Bulletin 214.