

to the meridian itself, will then be the north, or right line, and it equal distant from all parts of the horizon. Depend that point all observations out the horizon meridian to forty degrees (or more, as you have them) of the inhabitants in Madrid; and towards the point in Madrid comes to the meridian, you will find it in the north, or any of the globe, under forty degrees on minutes. Note: If it were required, to rectify the globe for South latitude, then you must elevate the south pole to the given latitude, instead of the north pole; but this new problem will be soon explained.

Prob. 3. The latitude and longitude of any place given, to find the time. You are first to observe whether the longitude be reckoned from London; or from the first meridian; for on some globes the first meridian begins twenty-three degrees, an exact twenty degree, and on some's globe, sometimes degrees west of London; but if you know where the first meridian is on the globe, it is very easy to know the difference from the meridian of London.

Prob. 4. To find the latitude of any place given, to tell the place that have the same latitude. Disposition: All those places that have the same latitude have the day and night of the same length, or the same length, at the same time of the year. Rule: Bring the given place, or place, to the meridian (suppose London fifty-one degrees thirty-nine minutes, and Madrid fifty degrees and minutes north); then turn the globe, and all those places that pass under forty degrees ten minutes have the same latitude as Madrid, which you will find to be: Peking nearly for one, and many other places.

Prob. 5. To tell the difference of the latitude of places. There are two variations, or rules.—First, If the latitude be both north or both south, then subtract the less from the greater latitude, and the remainder is the difference, or excess. Thus between London and Madrid is twelve degrees thirty-nine minutes, the first being fifty-one degrees thirty-nine minutes, and the other forty degrees. And between Candy and Stockholm is fifteen degrees thirty-nine minutes, for Stockholm is about fifty-one degrees thirty-nine minutes north, and Candy from degrees thirty-nine minutes north. Secondly, If one be to the north, and the other on the south-side of the equator, (thus is to be, if one be north and the other south latitude) then add them all together, and their sum is the difference of the latitude required.—Thus, Copenhagen is fifty-five degrees forty minutes north, and the Island of Madagascar is nineteen degrees thirty minutes south: these added together make seventy-four degrees ten minutes, the difference of longitude required in the problem.

Prob. 6. The longitude of any place given from any meridian, to tell those places, having the same longitude. This is done after the same manner as the other; only here the answer will be on the equator, in others upon the meridian. I would know what places have the same longitude as London, and the same longitude as Moscow? The rule is, bring London to the meridian, then all those places on the globe (from the north pole to the south part of the horizon) that lie under the edge of the meridian, have the same longitude as London: thus Port Norfolk, and Port Mores, in Guinea, have the same, or very nearly the same longitude as London.—And Moscow, in Moscow, has very nearly the same longitude as Aleppo, in Syria; also Scandinavia, Astrachan, and Tripoli, in Syria, have the same longitude, viz. thirty-seven degrees thirty minutes from London.

Prob. 7. To tell the difference of longitude of places. Rule: There are two variations. First, If the places lie both east, or both west, of the first meridian, or where you reckon the longitude from, viz. if they both be east, or both be west longitude, then subtract one from the other, you have the difference. Thus I find Jerusalem has thirty-four degrees fifteen minutes east longitude from London, and Peking one hundred and ten degrees fifty-two minutes east longitude; therefore subtract thirty-four degrees fifteen minutes from one hundred and ten degrees fifty-two minutes, and there remains seventy-four degrees thirty-seven minutes difference of longitude east or west; that is, Peking is seventy-four degrees, thirty-seven minutes east longitude of Jerusalem, or Jerusalem is seventy-four degrees thirty-seven minutes west longitude of Peking. Secondly, If one place be east, and the other west longitude of the first meridian, (suppose London, or any other meridian) then add their longitudes together, and the sum is the difference of longitude required.

Prob. 8. The day of the month given to find the sun's place in the ecliptic. Rule: The day of the month being given, look on the inner calendar on the new globe, and you have the sign, and the degree of that sign that the sun is in for that day, according to the new style. If it be upon old globe, look on the outer calendar, you have the sign, and degree of the sign. N.B. You may further observe, that the calendar used throughout Europe is the calendar for new style, viz. new style is always known from the other, because it has the solstis day, and several other things inscribed on the horizon.

Prob. 9. The sun's place given to find the day of the month. This is only the reverse of the former problem; for having the sun's place given, seek it in the circle among the signs, then against that degree in the calendar, new style, you have the day of the month required. Example: I would know what time of the year the sun is in one degree of Gemini; at also when he enters Aries, Cancer, Libra, and Capricorn. Proceed according to the rule, and you will find the day to be May 21, June 21, September 21, and December 21, in the last.

Prob. 10. The latitude and day of the month given, to find the sun's place in the ecliptic, and rectify the globe for it. Rule: Find the sun's place on the horizon by problem 8. and having noted what degree he is in, look upon the ecliptic on the globe, and find the same sign and degree as you did on the horizon; then bring that degree of the ecliptic very carefully to the graduated edge of the brass meridian, and holding the globe steady, turn the index exactly to the upper twelve (which represents twelve at noon), and thus is the globe rectified for that day, and the degree of the ecliptic that lies under the equator represents the sun's place at noon, or twelve o'clock that day. Note: The astronomer's day is reckoned from, or begins at, twelve o'clock; and if you fix the quadrant of altitude to the latitude in the month, the globe will be completely rectified.

Prob. 11. To tell the declination of the sun on any day of the year. Rule: Having found the sun's place in the ecliptic for the given day, bring it to the brass meridian, and observe what degree of the meridian it lies under, and whether it be on the north or on the south side of the equator, for that is the declination accordingly. Thus, on April 21, the sun has eleven degrees thirty minutes north declination, and on May 21, he has twenty degrees thirty minutes declination; but on October 27, he has twelve degrees thirty minutes south declination.

Prob. 12. A latitude and day of the month given, to tell the sun's meridian altitude, viz. his height at noon. Rule: Bring the sun's place to the meridian, and observe what degree of the meridian the sun's place is under: for those degrees on the meridian that are intercepted, or lie between the fourth verge of the horizon and the degree, which is over the sun's place on the meridian, (counted on the meridian) is the sun's meridian altitude required. Thus I find his meridian altitude at London, May the 21st, to be fifty-nine degrees, but on November the fifth, he has but twenty-three degrees thirty minutes altitude.

Prob. 13. The latitude and day of the month given, to tell the sun's altitude at any time. Example: On May the 21st, at nine in the morning, and at five in the afternoon, at London; I would know the sun's altitude, or height? Rule: Rectify the globe for the latitude, and bring the sun's place (one degree gemini) to the meridian, and the index to the upper twelve on the dial-plate; then screw the quadrant of altitude on the zenith, viz. the left edge of the quadrant be fixed on the meridian, at fifty-one degrees thirty minutes; then turn the globe till the index points to the hour, viz. nine in the morning; this done, fix the globe by thrusting a quill between it and the horizon; lastly, turn the quadrant about till the graduated or figured edge touch the sun's place, (viz. one degree gemini) and the degrees on the quadrant, counted from the horizon upwards on the quadrant, is his height at that time, viz. forty-three degrees thirty minutes; Then turn the globe till the index point at five in the afternoon; and also turn the quadrant on the west-side (without untwisting it) till it touches the sun's place; and you have about twenty-four degrees on the quadrant, his altitude at that time. N.B. At North cape, (north latitude, seventy-two degrees at nine in the morning) May the 21st, he will not be but about thirty-two degrees high.

Prob. 14. The latitude given, to tell the rising and setting of the sun, and length of the day and night, at any time of the year in any place. Rule: Rectify the globe, (viz. elevate it for the latitude; bring the sun's place to the meridian, and index to the upper twelve); then turn it till the sun's place comes even with, or lies right against the horizon verge, on the east-side of the horizon, then the index will show you the time of the sun's rising; turn it to the west-side, or verge of the horizon, and the index will show you the setting. Or thus: having got the hour the sun rises, count how many it wants of twelve, for so many hours will it set after. Thus: if the index points to four in the morning at rising, it will of course be at eight at night, &c. Proceed thus, and you will find the sun, on May the 26th, at London, to rise about five in the morning, and set at eight at night. Now double what he wants of twelve at rising, viz. eight hours, and is given the length of that day at London, viz. sixteen hours.

Prob. 15. To tell the sun's right ascension. Bring the sun's place to the brass meridian, and note what degree of the equator is cut by the meridian, for that is his right ascension required. I would know the sun's right ascension on March the 21st; June the 21st, September the 23d, and December the 21st? I find the sun's place for those different days, and bring it to the meridian; I find the meridian cut the equator in (0), in (90), in (180), and in (270 degrees) his right ascension required. Note: When the sun enters Aries, March the 21st, he has no right ascension, because it is counted from, or begins at Aries; therefore, on March the 20th, he must have his greatest right ascension, viz. three hundred and fifty-nine degrees.

Prob. 16. To find the sun's oblique ascension and descension at the time, and in any latitude. Rule: Rectify the globe for the latitude, and bring the sun's place down to the eastern verge of the horizon; then observe what degree the horizon cuts the equator in, for that is the oblique ascension required. Thus on March the 21st, June the 21st, September the 23d, and December the 21st, viz. when the sun enters Aries, Cancer, Libra, and Capricorn, you will find his oblique ascension at London to be (0), (56), (180), and (304). And on the same days his oblique descension will be (0), (123), (180), and (37) and a half.

Prob. 17. The latitude and day of the month given, to tell the sun's ascensional difference, viz. how much he rises, or sets, before and after day and consequently to tell the length of the days, supposing there were no index to the globe. Rule: By the last problem find the sun's right, and oblique ascension; then subtract the oblique from the right ascension, or the contrary, and the remainder is the ascensional difference required; which divide by fifteen, the degrees of the equator that pass through the meridian of one hour, (or seven and a half for half an hour) gives the answer in time, that the sun rises and sets before and after six. Thus, on May 26, I find the sun is five degrees of Gemini, and his right ascension is fifty-four degrees and on the same day, his oblique ascension is thirty-four degrees; now thirty-four degrees from fifty-four degrees there remains twenty degrees, his ascensional difference; which divide by fifteen gives two hours, the time that he rises before, or sets after six.

Prob. 18. The latitude and day of the month given, to tell the sun's amplitude, viz. his distance from the east and west points of the compass he rises and sets upon. Rule: The globe being rectified, bring the sun's place to the eastern verge of the horizon, (which shows his rising) then the degree upon the innermost circle of the horizon, counted from the true east point to the place where the sun's place lies against the horizon, shows you the sun's amplitude. Proceed according to the rule, you will find the sun's amplitude at London, (May 21,) at rising, to be about thirty-four degrees from the east to the north, and at setting thirty-four degrees from the west to the north; and the point he rises upon is north-east.